



# Computational evaluation of the bucky components influence on the estimation of normalized glandular dose in digital mammography

J. Godeli<sup>a,\*</sup>, D.M. Cunha<sup>b</sup>, M.P.A. Potiens<sup>c</sup>, M.E. Poletti<sup>a</sup>

<sup>a</sup> Departamento de Física, Faculdade de Filosofia, Ciências e Letras de Ribeirão Preto, Universidade de São Paulo, Ribeirão Preto, SP, Brazil

<sup>b</sup> Instituto de Física, Universidade Federal de Uberlândia, Uberlândia, MG, Brazil

<sup>c</sup> Instituto de Pesquisas Energéticas e Nucleares, Comissão Nacional de Energia Nuclear, IPEN/CNEN, São Paulo, SP, Brazil

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## ABSTRACT

The mean glandular dose (MGD) is the most suitable dosimetric quantity used in mammography to describe the absorbed dose by the breast, although it cannot be directly acquired. Studies have provided conversion factors widely implemented in international dosimetry protocols to estimate MGD, such as normalized glandular dose (DgN). Over time, the DgN estimation was refined by considering geometric models that approach a real clinical environment, such as new anode/filter combinations, compression plate and breast models. However, there is no detailed study of how the bucky (support plate, antiscatter grid and detector) can affect the DgN estimation. A modified PENELOPE Monte Carlo code was used for DgN estimation. The irradiation geometric model was built as a complete digital mammography system, considering a homogeneous breast and different typical bucky models in commercial mammography units. Simulations were carried out for mono and polyenergetic beams considering different imaging geometries. Studies with monoenergetic beams showed that the bucky presence affected DgN mainly for higher beam energies and thinner breasts. The breast support plate was the bucky component that most affected the DgN, followed by the anti-scatter grid and finally, the image detector. Studies with polyenergetic conventional (low-energy) spectra showed that the bucky exerted a minimal influence on DgN values (less than 1.0%). For high-energy spectra, mainly employed in modalities such as contrast-enhanced digital mammography, the DgN values were more affected by the bucky, increasing by 4.8% the DgN values for a 2 cm thick breast and a W/Cu 50 kV spectrum. Bucky inclusion in computer simulations is highly recommended mainly for thinner breasts and high-energy spectra. To simplify the simulations, we confirm that a homogeneous carbon fibre block support, with thickness between 3.9 and 4.1 mm, can be used as a good substitute for a complete bucky model.

## 1. Introduction

The International Agency for Research on Cancer (IARC) and American Cancer Society points to breast cancer as the leading cause of cancer deaths among women, with more than 680.000 reported deaths worldwide in 2020 (Sung et al., 2021). Mammography imaging technique is currently the most suitable method for the diagnosis and prevention of cancer, although this imaging modality presents risks of radiation-induced carcinoma (Miglioretti et al., 2016). To measure this risk, Karlsson et al. (1976) proposed the estimation of the mean glandular dose (MGD). This quantity is currently the most recommended test for quality control in mammography (Dance and Sechopoulos, 2016) and is defined as the product between the entrance skin air kerma (ESAK) or exposure and a conversion factor, called by several authors as normalized glandular dose (DgN) (Stanton et al., 1984; Wu et al.,

1994; Boone, 2002; Fedon et al., 2015; Nosratieh et al., 2016; Sarno et al., 2017).

Initially, simple geometric models representing clinical conditions were employed to reproduce a complete mammography environment, due to computational limitations. Dance (1990) estimated the normalized glandular dose as a function of the half value layer (HVL), considering only the homogeneous breast surrounded by a skin layer, the breast compressor and an image receptor. Subsequently, other studies have extended DgN estimates to other conventional (Wu et al., 1991, 1994; Boone, 1999; Dance et al., 1999, 2000, 2009) and high-energy X-ray beams (used in imaging techniques such as contrast-enhanced, tomosynthesis and dual-energy) (Dance, 1990; Boone, 1999; Zhang et al., 2012; Cunha et al., 2013; Dance and Young, 2014; Nosratieh et al., 2016). Further works observed the influences of the spectra

\* Corresponding author.

E-mail address: [juliodeli@usp.br](mailto:juliodeli@usp.br) (J. Godeli).

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characteristics, such as tube potential, anode/filter combination and source-to-skin distance (Wu et al., 1991, 1994; Boone, 1999; Dance et al., 2000). Dance et al. (2000) proposed the *g-factor* as a correction for different anode/filter combinations, showing that the DgN values can vary up to 6.1% for a fixed tube potential. Hemdal (2011) pointed out the importance of the compression plate in glandular dose estimations. The influence of different breast models was also analysed, varying their elemental composition (Carlsson and Dance, 1992), glandular fraction, distribution and positioning of glandular tissue (Wu et al., 1991, 1994; Boone, 1999; Dance et al., 2000, 2005; Hernandez et al., 2015) and skin models (Huang et al., 2008; Sarno et al., 2017; Massera and Tomal, 2018). Breast dosimetry protocols were established in order to standardize the conversion factors. The European (Engen et al., 2013), IAEA (IAEA, 2007) and United Kingdom (IPEM, 2005) protocols are based on the conversion factors proposed by Dance and collaborators (Dance, 1990; Dance et al., 2000, 2009; Dance and Young, 2014; Dance and Sechopoulos, 2016). Over the years, Dance and collaborators expanded the conversion factors for a wide range of spectra, with different anode/filter combinations and tube potentials. Previously, the American protocol (ACR, 1999) used a method adapted from Wu et al. (1991, 1994), using a standard ACR phantom. This protocol, published in 1999, is outdated as it was developed solely for screen-film mammography. In 2018, a new ACR mammography quality control manual was published (ACR, 2018), expanding its data for digital mammography and breast tomosynthesis. This manual uses a different method, proposed by Dance and collaborators (Dance et al., 2000, 2009, 2011). Although computational simulations evolved and become more complete over the years (Sechopoulos et al., 2023), there is no detailed information on the literature on how the mammography bucky can influence DgN estimations.

The mammography bucky can be simplified as a set of three components: First, the breast support plate, which can be considered as a homogeneous plate composed of a low atomic number material, such as polymethylmethacrylate (PMMA) or carbon fibre (Dance et al., 2000; Sechopoulos et al., 2007; Massera and Tomal, 2018). Second, the antiscatter grid, containing strips with a high atomic number material, such as lead or copper, separated by an interstitial space composed of materials that allow a high photon transmission, such as PMMA or air (Dance et al., 1992; Rezendes et al., 1999; Cunha et al., 2012). Third, the image detector, which consists of a thin photosensitive material, such as caesium iodide (CsI) or amorphous selenium (aSe) (Yaffe, 2010).

This work presents a detailed study on how the mammography bucky can affect the estimate of DgN values and evaluates the bucky influence on DgN values previously tabulated on international protocols, which use homogeneous breast models. For this, computer simulations were performed by considering a complete geometric model built on a modified Monte-Carlo PENELOPE code (Salvat et al., 2009). Initially, several tests were carried out with monoenergetic beams, in order to study the direct influence of each bucky component on DgN values for different conditions, such as beam energy, configuration of each bucky component and breast thickness. Later, studies with polyenergetic beams were carried out. For this step, geometric models and simulation conditions previously established by the literature were reproduced, with and without the bucky presence, in order to evaluate how the bucky inclusion can influence the pre established DgN values in literature, on which the international mammography protocols are based on. Additionally, studies with homogeneous blocks with different thicknesses placed under the breast were carried out in order to evaluate a suitable substitute for the whole bucky simulation.

## 2. Materials and methods

### 2.1. Study of the bucky influence on monoenergetic values of DgN

#### 2.1.1. Geometric model

The simulated geometric model was based on recommendations of the report AAPM TG-195 case III (Sechopoulos et al., 2015). Some alterations were made in order to consider a digital mammography system

with all bucky components included, as shown on Fig. 1. The model includes an X-ray source, a chest wall and a compressed semicircular breast, resting on a bucky represented by a support plate, an antiscatter grid and an image detector.

All bucky components were modelled with the PENELOPE package (Almansa et al., 2021) of the PENELOPE Code, defined by means of their limiting quadratic surface. The X-ray source with a focal spot of  $0.3 \times 0.3 \text{ mm}^2$  was positioned 66 cm above the detector. A rectangular ( $30 \times 30 \times 170 \text{ cm}^3$ ) block of water represented the chest wall. The compression paddle was simulated as a PMMA block, with the dimensions similar to the detector ( $18 \times 24 \text{ cm}^2$ ) and thickness of 3 mm (Cunha et al., 2012; Del Lama et al., 2017). The breast was modelled as an 8 cm radius semi-cylinder, with thickness from 2 to 8 cm. Its inner layer was composed by a homogeneous mixture of adipose and glandular tissues, surrounded by an outer layer of 1.45 mm thick skin tissue (Huang et al., 2008). The glandular fraction was fixed at 50%. The elemental compositions and densities of breast and skin tissues were based on the data provided by Hammerstein et al. (1979). The bucky was separated into three components: breast support plate (BSP), antiscatter grid (ASG) and image detector (ID). Several models of these components were considered, according to the literature (Dance et al., 1992; Rezendes et al., 1999; Massera and Tomal, 2018; Dance et al., 2000; Zhao and Zhao, 2008; Dance et al., 2009; Yaffe, 2010; Cunha et al., 2012). A cover was considered for the antiscatter grid, composed of 0.3 mm carbon fibre plus 0.13 mm polyester (Dance et al., 1992), placed on both extremities of the grid. The detectors were modelled as homogeneous plates. Table 1 shows each bucky component considered in this work. The combinations of bucky components are also described on Table 1, being based on bucky configurations present on real mammography systems.

#### 2.1.2. Monte Carlo simulations and score quantities

The simulations were carried out using the PENELOPE code (version 2008) (Salvat et al., 2009). To obtain good accuracy in the results, the simulations were performed with  $10^8$  primary photons. Each run took from 150 to 600 min (processor Intel Core i7-3770k CPU @ 3.50 GHz). The minimal statistical uncertainty obtained, as proposed in AAPM TG-195, was approximately 0.1%. Monoenergetic photons from 5 to 50 keV were considered. The simulations were carried out considering an irradiation field that covered all the detector surface ( $18 \times 24 \text{ cm}^2$ ). Both photons and electrons were followed, tracking their path until the cut-off energy of 1 keV. The electrons were tracked down since they can exert influence over fluorescent-generated photons from materials with medium and high atomic numbers.

The PENELOPE code was adapted to compute the energy ( $E_i$ ) and incident angle ( $\theta_i$ ) of photons that reached a region of interest (A), positioned directly under the compression paddle, with area of  $30 \times 30 \text{ mm}^2$  and placed 40 mm from the chest wall (disregarding backscatter) (Dance and Young, 2014). Thus, the incident Air Kerma  $K_{air}$  was calculated, in mGy, through Eq. (1) (Sarno et al., 2017).

$$K_{air} = \sum_i \frac{E_i}{A \times \cos(\theta_i)} \times \left( \frac{\mu_{tr}}{\rho} (E_i) \right)_{air} \quad (1)$$

where  $\left( \frac{\mu_{tr}}{\rho} \right)_{air}$  is the air mass energy transfer coefficient. In order to simplify our simulations, we implemented analytical exponential equations on the main program, fitted to calculate the  $\left( \frac{\mu_{tr}}{\rho} \right)_{air}$ , in  $\text{cm}^2/\text{g}$ , based on ICRU-90 database (ICRU, 2016) considering the energies between 1 and 50 keV. The fitted equations were validated by comparing our calculations with the NIST database (Higgins et al., 1991), observing very good agreement (difference lesser than 1%) between our results and the literature.

The mean glandular dose was calculated, in mGy, by Eq. (2):

$$MGD = \sum_i \left( \frac{E_{abs_i}}{M \times g_r} \right) \times G \quad (2)$$

where  $E_{abs_i}$  is the energy absorbed by the breast after each photon interaction,  $M$  is the breast tissue mass (without skin tissue),  $g_r$  is the

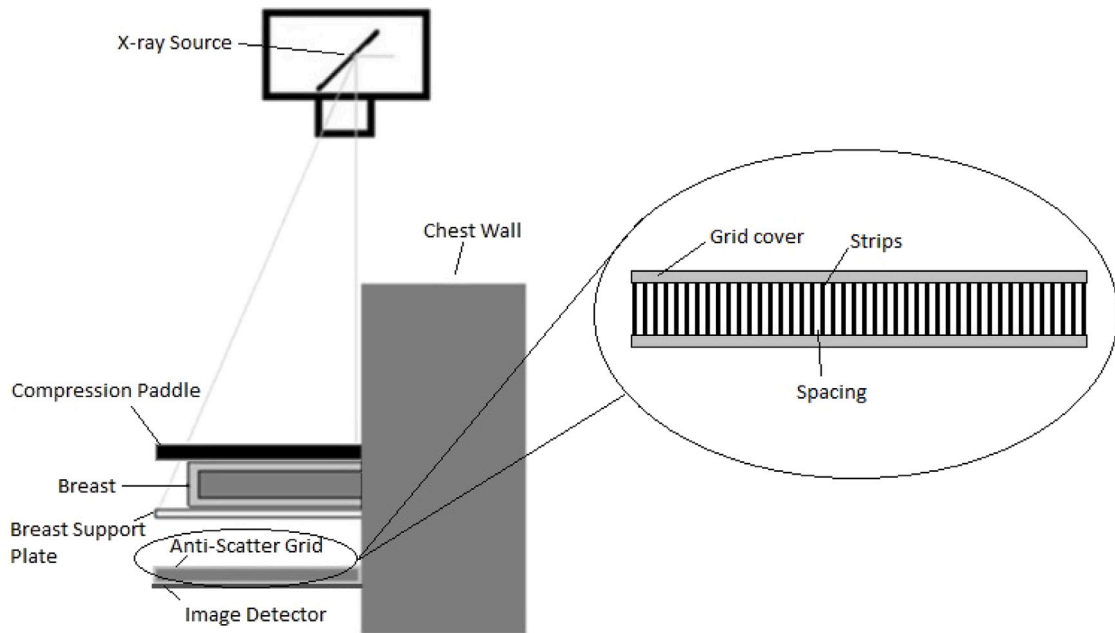


Fig. 1. Simulated geometric model (not to scale).

**Table 1**  
Bucky components and its combinations considered in this work.

Breast Support Plate (BSP)				
Material	Density (g/cm <sup>2</sup> )	Thickness (mm)	Literature	
Carbon Fibre	1.45	1.20	Dance et al., 2000	
Carbon Fibre	1.45	2.10	This Work	
Carbon Fibre	1.45	4.10	Dance et al., 2009	
PMMA	1.19	2.00	Massera and Tomal., 2018	
antiscatter Grid (ASG)				
Grid Strips	Interspace	Height (mm)	Frequency (cm-1)	Literature
Pb (Linear)	PMMA (0.3 mm)	1.50	31 lines	Dance et al., 1992
Cu (Cellular)	Air (0.64 mm)	2.40	15 cells	Rezentes et al., 1999
Image Detector (ID)				
Material	Thickness (μm)	Literature		
CsI	150	Yaffe, 2010 ; Cunha et al., 2012		
aSe	250	Zhao and Zhao, 2008 ; Yaffe, 2010		
Complete Bucky (CB)				
Combinations				Label
2.1 mm carbon fibre BSP + Linear ASG (Pb+PMMA) + CsI ID				CB1
2 mm PMMA BSP + Cellular ASG (Cu+Air) + aSe ID				CB2
2.1 mm carbon fibre BSP + aSe ID				CB3

breast mass glandular fraction and  $G$  is a weighting factor, proposed by Boone (1999), that estimates the energy fraction absorbed by the glandular tissue:

$$G = \frac{g_r \cdot \left(\frac{\mu_{en}}{\rho}\right)_{gl}}{g_r \cdot \left(\frac{\mu_{en}}{\rho}\right)_{gl} + (1 - g_r) \cdot \left(\frac{\mu_{en}}{\rho}\right)_a} \quad (3)$$

where  $\frac{\mu_{en}}{\rho}$ , is the mass energy absorption coefficient of glandular ( $gl$ ) and adipose ( $a$ ) tissues. These coefficients were estimated based on their elemental compositions (Hammerstein et al., 1979) and according the mixture rule, with values for mass energy-transfer coefficients provided by NIST (Higgins et al., 1991). These calculations were validated by comparing our values for  $\frac{\mu_{en}}{\rho}$  with those provided by NIST (Higgins et al., 1991), for the adipose tissue. We observed a very good agreement (differences lesser than 1%) between our values and the literature, within the energy range considered in this work. The  $G$ -factor was estimated interaction-by-interaction (Wilkinson and Heggie, 2001).

The normalized glandular dose (DgN) was estimated, in mGy/mGy, by the following equation (Stanton et al., 1984; Wu et al., 1991; Boone, 1999):

$$D_g N = \frac{MGD}{K_{air}} \quad (4)$$

### 2.1.3. Simulation steps

To assess the bucky influence on DgN values as a function of the beam energy and breast thickness, simulations were performed including each bucky component step-by-step, (i) first simulating only the breast (considered as a reference), (ii) later adding the breast support plate, followed by (iii) the antiscatter grid and finally, (iv) the image detector. Simulation of the grids were conducted separately from the rest of the bucky components, due to their complex geometry, which requires a large number of objects to simulate the grid strips or cells. For the grid geometry, the grid strips were reproduced with the CLONE function of the PENGEOM (Almansa et al., 2021). Simulations were performed with photons with energies between 1 and 50 keV and incident angles between 0 and 89° from the  $z$ -axis. In order to score the transmitted, backscattered and fluorescence photons in function of the photon energy and incident angle, impact detectors were placed on top and bottom of the grid. Were assessed the characteristic photons from the radiopaque material and the overall backscatter from the whole grid, regardless of which component (cover, interstitial space and radiopaque material) the photon interacted with. The method was validated through comparison with Zhou et al. (2016) data of transmission ratios of primary and scattered photons, adapted to the

authors conditions, considering a 4 cm PMMA phantom placed above a linear grid, irradiated by polyenergetic beams (Rh-0.27 mm Cu). Our data showed very good agreement with the literature (differences lesser than 3%).

These parameters were then used as input in the main simulations of the bucky to estimate the ASG influence on DgN. As the tracked photon reached the antiscatter grid position, a series of Monte Carlo assessments were conducted. The probability of the photon transmission was estimated, according to its energy and incident angle. If the photon was transmitted, its tracking continued normally. On the other hand, if the photon interacted with the grid, the probability of backscatter or fluorescence were estimated.

For the tests analysing the image detector influence on DgN values, we considered the probability of transmission through the grid as well. Photons that backscatter from the detector surface and fluorescent-generated photons that reached the antiscatter grid position also had its transmission evaluated.

The relative influences of each bucky component on the DgN values were briefly analysed, varying their models and composition referenced in literature (Dance et al., 1992; Rezendes et al., 1999; Massera and Tomal, 2018; Dance et al., 2000; Zhao and Zhao, 2008; Dance et al., 2009; Yaffe, 2010; Cunha et al., 2012).

## 2.2. Comparison between polyenergetic DgN values acquired with and without the bucky presence using geometric models and simulation conditions provided by literature

To evaluate the bucky influence on DgN values tabulated on European breast dosimetry protocols (Engen et al., 2013; IPEM, 2005; IAEA, 2007), the geometric models and simulation conditions developed by Dance (1990) and Dance and Young (2014) were reproduced. The geometric model adopted by Dance (1990) consists of a semi-cylindrical breast with 8 cm radius and thickness ranging from 2 to 8 cm, surrounded by a 5 mm thick layer of adipose tissue, representing the skin. Dance and Young (2014) adopted a geometric model similar to that presented by Dance (1990), however, considering a homogeneous plate composed of 4.1 mm thick carbon fibre, positioned under the breast.

DgN values tabulated in American protocols (ACR, 1999, 2018) and vastly referenced in literature (Sarno et al., 2017; Ding and Molloy, 2017; Fedon et al., 2019) were also evaluated. For this step, the geometric model and simulation conditions proposed by Wu et al. (1994) was reproduced and studied. The geometric model adopted by Wu et al. (1994) consists of a 3 cm thick semi-elliptical breast ( $r_a = 9$  cm and  $r_b = 8$  cm) with a 4 mm thick layer of skin.

Dance (1990), Dance and Young (2014) and Wu et al. (1994) used their own codes to perform the simulations.

For the estimation of DgN using the geometric models of Dance (1990) Dance and Young (2014), the formalism adopted by Dance et al. (2000) was used, where it is considered:  $DgN = g.c.s$ . Where,  $g$  is the conversion factor of air Kerma to MGD, considering a breast with 50% glandular fraction,  $c$  is the correction factor for other breast glandularities and  $s$  is the correction factor for the spectrum considered.

The influence of all bucky models on the DgN values, described in Table 1, were analysed through simulations carried out with the geometric model proposed by Dance and Young (2014).

The simulations were performed using conventional and high-energy mammography X-ray beams, consistent with those used by each author. X-ray spectra were obtained for different anode/filter combinations and tube potentials through the model proposed by Boone et al. (1997). Table 2 specifies the spectra used in each geometric model.

The simulations were carried out with and without the inclusion of the bucky, being considered relevant any increase in DgN values greater than the measured uncertainty of 0.1%.

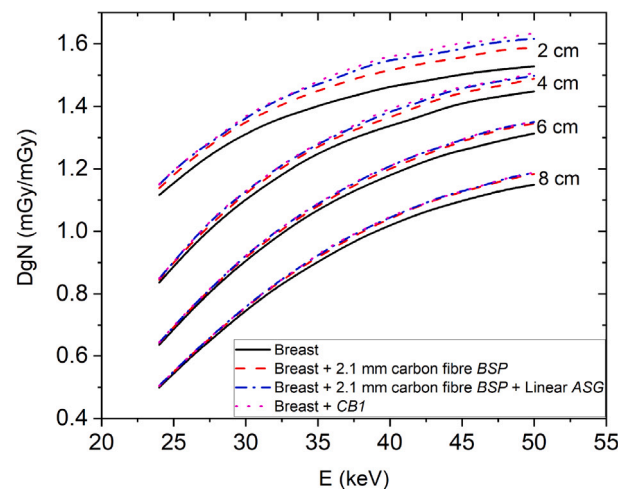


Fig. 2. DgN values as a function of the beam energy, for breasts with thicknesses varying from 2 to 8 cm, considering a addition of each bucky component.

## 2.3. Validation

The modification of PENELOPE code for calculating MGD values was validated through the report AAPM TG-195 case III recommendations (Sechopoulos et al., 2015). For comparison, the results of Sarno et al. (2017) were also presented alongside those found in the report (AAPM, 2015). Using the monoenergetic beam, the observed differences between our results and those of AAPM-TG195 and Sarno et al. were 0.25% and 1.11%, respectively. For the polyenergetic beam, the difference were 0.33% and 0.71%, respectively. A good agreement was observed between the simulated results in this work and the literature for both monoenergetic and polyenergetic beams, which validates the adaptation of the PENELOPE code for MGD calculations.

## 3. Results and discussions

### 3.1. Study of the bucky influence on monoenergetic values of DgN

#### 3.1.1. Overview of the whole bucky influence

Fig. 2 shows the computed DgN values, considering a 50% glandular breast model covered by 1.45 mm of skin, positioned upon a bucky corresponding to the CB1 configuration, as shown in Table 1. Simulations carried out with photons with energies up to 24 keV showed that the bucky inclusion had a minimal influence on the DgN values. Therefore, these results are not shown in this step.

Fig. 2 illustrates the influence of the inclusion of each bucky component on simulated DgN values, considering 2 to 8 cm thick breasts. From the figure it is observed that when the whole bucky is considered, the DgN values increase mainly for higher energies and thinner breasts. The 2 cm thick breast was the most affected by the bucky presence, showing an increase of 7% in DgN value for 50 keV beam. For this energy, the 4, 6 and 8 cm thick breasts presented an increase of 4.3%, 3.3% and 3.1% in DgN values, respectively. This reduction in the bucky influence for thicker breasts can be explained by the decrease in photons transmitted through the breast, affecting the relative contribution of the bucky region directly below the breast surface.

Additional tests were conducted with different breast models, varying their glandular fractions. Although the same trends as those presented for the 50% glandular breast were found, there was a minimal reduction on the bucky influence over the DgN with the increasing glandular fraction.

Further test were carried for a fixed 50% glandular breast and varying the skin tissue model (Massera and Tomal, 2018). The same trends and no significant differences on the bucky influence over the DgN values for all the skin models tested were observed.

**Table 2**  
X-ray spectra used in this work to reproduce DgN values obtained in the literature.

Conventional mammographic spectra			
Anode/Filter	Tube potential (kV)	HVL range (mmAl)	Literature
Mo - 0.03mmMo + Al	20 - 35	0.25 - 0.45	Dance, 1990
Mo - Rh	35	0.44 - 0.54	Wu et al., 1994
High-energy mammographic spectra			
Anode/Filter	Tube potential (kV)	HVL range (mmAl)	Literature
W - Al	40 - 50	1.00 - 2.00	Dance, 1990
W - 0.3mmCu	40 - 50	2.57 - 3.54	Dance and Young, 2014

**Table 3**  
Relative contribution of each bucky component over DgN values. For ASG and ID, the proportion of fluorescent energy in each contribution is shown in parentheses.

Relative contribution per component (%) - CB1			
Energy (keV)	BSP - 2.1 mm Carbon Fibre	Linear ASG (% Fluorescence)	CsI ID (% Fluorescence)
24	65.8	33.7 (52.3)	0.5 (0.00)
34	62.6	23.9 (24.2)	13.5 (97.5)
42	60.3	23.8 (11.6)	15.9 (99.1)
50	62.0	26.3 (6.1)	11.2 (98.3)
Relative contribution per component (%) - CB2			
Energy (keV)	BSP - 2 mm PMMA	Cellular ASG (% Fluorescence)	aSe ID (% Fluorescence)
24	75.7	21.8 (11.0)	2.5 (91.5)
34	75.4	23.3 (3.3)	1.3 (68.6)
42	71.0	28.0 (1.4)	1.0 (42.6)
50	67.5	31.6 (0.0)	0.9 (24.1)
Relative contribution per component (%) - CB3			
Energy (keV)	BSP - 2.1 mm Carbon Fibre	aSe ID (% Fluorescence)	
24	63.0	37.0 (93.8)	
34	78.0	22.0 (76.2)	
42	83.0	17.0 (52.4)	
50	83.8	16.2 (32.1)	

3.1.2. Detailed studies of the relative influence of each bucky component

To better illustrate our results, Table 3 shows the relative contribution (alongside its fluorescent energy absorption proportion) of each bucky component for the DgN values increase. Our studies shown that the contributions of each component are similar for all breast thickness considered in this work.

Our results show that breast support plate was the most influential bucky component over the DgN values increment, for all bucky models and every energy considered. This dominance can be explained by the high intensity of backscattered photons from the breast support plate that directly reach the breast surface. In addition, this component serves as an attenuating barrier for photons backscattered by the antiscatter grid and the image detector.

For a better analysis, Fig. 3 shows the influence over the DgN values with the consideration of various BSP models, as described in Table 1. For comparison, the influence exerted by the complete bucky (CB1) over the DgN is shown alongside the results. In this analysis, was considered a 2 cm thick breast.

Considering the carbon fibre BSP with thickness from 1.2 to 4.1 mm (Dance et al., 2000, 2009; Dance and Young, 2014), it was observed that DgN values increase linearly with thickness. This can be explained by the increase of the backscattering for thicker breast support plates. It is also observed that the DgN values estimated with the CB1 are close to those acquired for the 4.1 mm thick BSP. These results are consistent with the observations of Dance et al. (2009), Dance and Young (2014) and will be further explored in Section 3.2. Comparing the results for the 2 mm PMMA (Massera and Tomal, 2018) and a 2 mm carbon fibre support plate (acquired through linear interpolation between 1.2 and 2.1 mm carbon fibre BSP), our results show that the carbon fibre support has a slightly higher influence on the DgN values. This little difference was expected due to the fact that the carbon fibre has a higher density than the PMMA, providing a slightly higher number of interactions suffered in this type of support.

The second most influential bucky component on DgN values was the antiscatter grid. The contribution of the ASG over the DgN values can be separated in two main sources: the characteristic X-rays from the

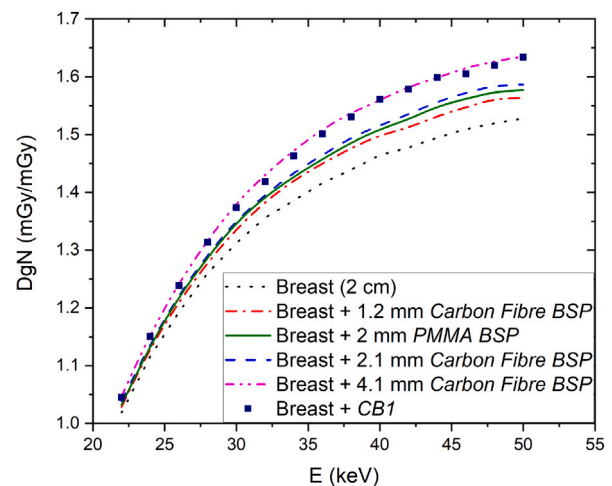


Fig. 3. DgN values as a function of the beam energy for breasts a 2 cm thick breast for the BSP models shown on Table 1 and the CB1 inclusion.

radiopaque strips or cells and the backscattering from the interstitial space and the grid cover.

First observing the linear grid (Pb/PMMA) present on CB1 shown in Table 3, it can be observed that for lower energies, the  $L_{\alpha}$  ( $\approx 10.54$  keV) and  $L_{\beta}$  ( $\approx 12.62$  keV) fluorescent photons from the Pb strips (Sánchez Del Río et al., 2003) are the main source of the relative contribution of the linear ASG on DgN values due to its high production probability at lower energy, rapidly decreasing for higher beam energies. As energy increases, the backscattering from the interstitial space and the grid cover became more influential, mostly due to the increase likelihood of occurring backscattering within the PMMA interstitial space. For all energies, there was no relevant contribution on DgN values from the backscattering on the lead strips.

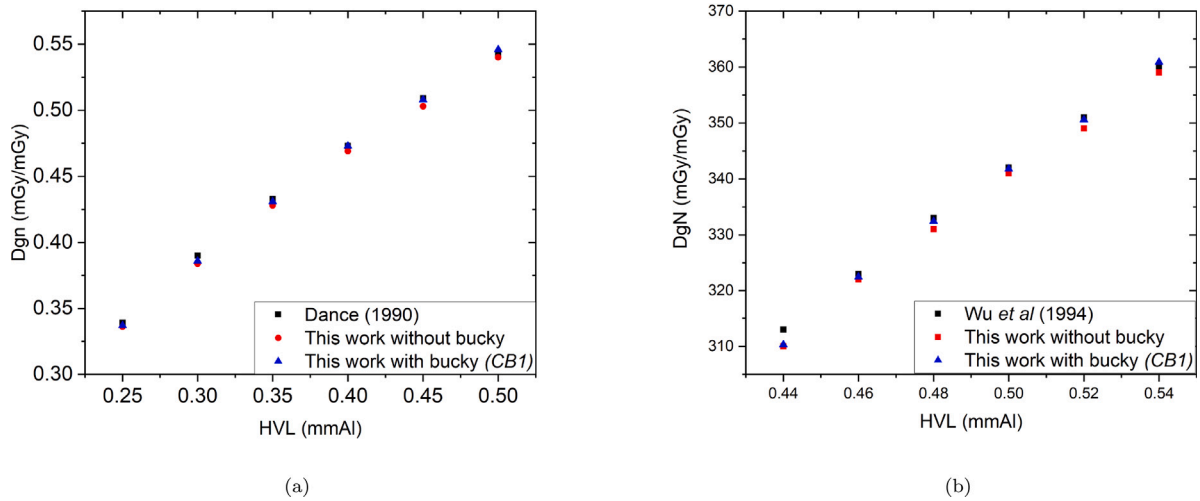


Fig. 4. Comparison between DgN values as a function of half-value layer (HVL) for the geometric model and simulation conditions proposed by Dance (1990) (2 cm thick breast) and Wu et al. (1994) (3 cm thick breast), with and without the inclusion of a bucky (CB1). In addition, the DgN values obtained by the authors are presented.

Now observing the cellular grid (Cu/air) present on CB2 shown in Table 3, the Cu fluorescence photons from the cells exerted little contribution to the overall antiscatter grid influence on DgN values. This was expected due to the relatively low energy of the copper  $K_{\alpha}$  ( $\approx 8.04$  keV) and  $K_{\beta}$  ( $\approx 8.91$  keV) fluorescent photons (Sánchez Del Río et al., 2003). This low energy is highly attenuated on the way to the breast tissue, and become less likely to occur as the energy increases. Although in this case, it was observed that the copper strips have a increase of backscattering for higher energies. For all energies, the contribution from the interstitial space (air) backscattering was irrelevant.

The least influential component was the image detector for all energies considered and systems simulated. For both detectors present on CB1 and CB2, this low influence was expected due to the presence of the ASG that blocks the majority of the emerging photons from the detector. From Table 3, it is observed that the CsI detector only exerted a little contribution on the CB1 overall influence for energies above 34 keV, due the characteristic X-rays from the detector. Now observing the aSe detector present on CB2, it is observed that its influence is negligible for all energies. This little influence decreases even more with the increasing energy, due the decrease of the probability of fluorescence photons to occur. Tests carried out without the antiscattering grid (CB3) showed that the relative contribution of the aSe detector greatly increases, reaching 37% of the total influence on DgN increment for 24 keV beam. However as seen previously, the detector contribution decreases for higher energies, due the decrease of the fluorescence production.

### 3.2. Studies carried out with geometric models and simulation conditions provided by literature

#### 3.2.1. The minimal influence of bucky inclusion on DgN values for studies with conventional beams

Fig. 4 shows the simulated DgN values for conditions and geometric models proposed by Dance (1990) and Wu et al. (1994) for a 2 and 3 cm thick breast, respectively.

For both comparisons shown in Fig. 4, the bucky inclusion on the geometric model exerted minimal influence over DgN values (less than 1.0%). Nevertheless, a slight discrepancy between the results with and without the bucky inclusion is observed for higher HVL. This is expected due to the increase of the mean energy of the beam and it is an indicator of how the bucky presence can affect the DgN values when high-energy spectra is considered.

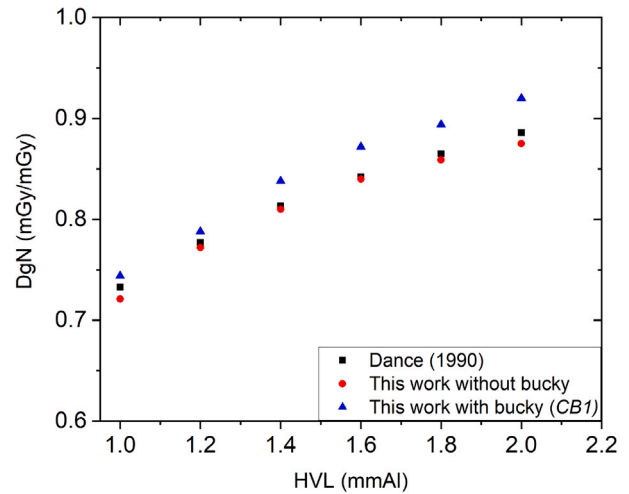


Fig. 5. Simulated DgN values in this work with and without the inclusion of a complete bucky (CB1) model, using the methodologies and geometric models proposed by Dance (1990), for a 2 cm thick breast with 50% glandular fraction.

#### 3.2.2. The importance of the bucky consideration for studies with high-energy beams

Fig. 5 shows the simulated DgN values for the conditions and geometric models proposed by Dance (1990), with and without the inclusion of a bucky (CB1), followed by the literature results. The simulations were performed using the high-energy mammography spectra (W/Al) described in Table 2. A 2 cm thick breasts was considered.

Fig. 5 shows the bucky influence on DgN values increasing with HVL. For HVLs of 1.0 and 2.2 mmAl, the bucky inclusion increased the DgN values by 3.1 and 5.1%, respectively. High-energy beams have greater transmission ratio in thinner breasts. In addition, due to the higher energies that compose these spectra, it is expected that a larger number of photons backscatter on the bucky, causing the observed increase of their influence on DgN values in comparison to conventional (low-energy) beams.

To analyse the influence of kVp over the bucky influence on the DgN, additional tests were carried out using various spectra (W/Al) with tube potentials between 40 and 50 kV and fixed HVL at 2.2 mmAl. The bucky influence on DgN values remained practically unchanged, varying up to 0.2%. This observation shows that the bucky influence

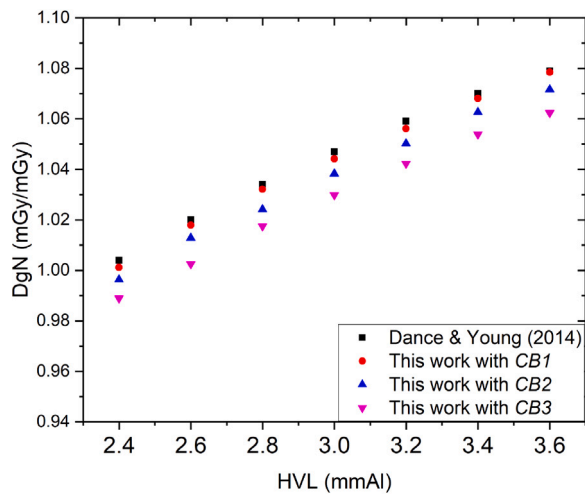


Fig. 6. Simulated DgN values for the methodology proposed by Dance and Young (2014), accurately reproducing its geometric model with a 4.1 mm carbon fibre, or using bucky models proposed in Table 1. DgN values were computed for a 4 cm thick breasts with 50% glandular fraction.

over the DgN depends strongly on spectra HVL and mean energy, but minimally on kVp.

Since the simulation of a complete bucky can be complex and time demanding, a solution would be the substitution of the whole components for a homogeneous carbon fibre block placed right under the breast, as previously suggested by Dance et al. (2009).

### 3.2.3. The variation of the influence of different bucky models on DgN values and the analysis of the consideration of a homogeneous block of carbon fibre as a substitute for simulating the complete bucky

Fig. 6 shows the simulated DgN values for the conditions and geometric models proposed by Dance and Young (2014), with the inclusion of all three complete models specified in Table 1, followed by the literature results. The simulations were performed using the high-energy mammography spectra (W/Cu) described on Table 2. A 4 cm thick breast was considered. The geometric model proposed by Dance and Young (2014) considers a 4.1 mm thick carbon fibre support. For comparison purposes, this component was replaced by the complete bucky models (CB1, CB2 and CB3), described on Table 1. Observing Fig. 6, it is noticed that the DgN values acquired with CB1 are in good agreement with the literature, showing similar results. However, simulations with CB2 and CB3 which have less influential components (cellular grid for only CB2 and aSe detector), presented smaller DgN values (up to 1.5%) if compared with the literature.

Our tests showed that consideration of a 4.1 mm thick carbon fibre block in simulations can replace a complete bucky model similar to CB1, in agreement with the literature (Dance et al., 2009; Dance and Young, 2014). Although the literature recommendation can overestimate the influence of other bucky models. Additional tests showed that the bucky models CB2 and CB3 can be accurately replaced by a 4 mm and 3.9 mm thick carbon fibre blocks, respectively.

## 4. Conclusion

The tests carried out with monoenergetic beams showed that implementation of the bucky in the geometric model can be relevant for DgN acquisitions that considers thinner breasts and X-ray energies higher than 24 keV. The DgN was more influenced by the breast support plate for all considered geometric models, being responsible by up to 60% of the total increase of DgN values, depending on which model was adopted. The antiscatter grid showed a minor, although noticeable influence on the DgN increment. The image detector showed a small

importance on the results, being relevant only for geometric models that do not consider the antiscatter grid.

Tests carried out with polyenergetic beams showed that the consideration of the bucky on the acquisition of DgN is highly recommended when high-energy mammography X-ray spectra is considered, which can be relevant on the application of advanced mammographic techniques, such as contrast-enhanced mammography. To reduce the time of simulation of DgN values and simplify the geometric models considered, it was observed that the whole bucky can be substituted by a homogeneous carbon fibre block with thickness between 3.9 and 4.1 mm, depending on its components.

Many studies have shown that the energy deposition in glandular tissues can change with its distribution (Oliver and Thomson, 2019) and how the MGD and DgN can drastically change if compared with homogeneous breast models (Hernandez et al., 2015; Chang et al., 2020; Ferrauche et al., 2023; Sechopoulos et al., 2023). Therefore, there is a necessity to expand the studies towards heterogeneous breast models.

## CRedit authorship contribution statement

J. Godeli: Writing – original draft, Visualization, Validation, Project administration, Methodology, Investigation, Formal analysis, Data curation, Conceptualization. D.M. Cunha: Writing – review & editing, Validation, Investigation, Data curation. M.P.A. Potiens: Resources, Project administration, Funding acquisition. M.E. Poletti: Writing – review & editing, Visualization, Validation, Supervision, Resources, Project administration, Investigation, Funding acquisition, Conceptualization.

## Declaration of competing interest

The authors declare the following financial interests/personal relationships which may be considered as potential competing interests: Julio Godeli Neto reports financial support was provided by Coordination of Higher Education Personnel Improvement. Martin Eduardo Poletti reports financial support was provided by National Council for Scientific and Technological Development. Maria da Penha Albuquerque Potiens reports financial support was provided by State of Sao Paulo Research Foundation. If there are other authors, they declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.

## Data availability

Data will be made available on request.

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